

Supplementary information

Coseismic folding during ramp failure in the Sulaiman thrust belt

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Introduction

This supporting information provides a detailed description of the methodology used for InSAR data processing and Bayesian inversion (Text S1), supported by Figures S1 to S6, and Tables S1 to S3. Text S2, Figures S7, S8, and Tables S3 to S6 support the main text of section 2, describing geometry of the kinematic inversions of FBF, and FPF models and determination of the distribution of the coseismic slip. Text S3 defines the potency density, in support of section 2 of the main text.

Text S1

1 Methods of InSAR processing and modeling

1.1 InSAR analysis

We select two pairs of Sentinel-1A observations of ascending track $T - 071$ and descending track $T - 005$ in TOPS mode that have the shortest perpendicular and temporal baselines with the least seasonal atmospheric variation to retain high correlation. Features of these datasets including orbit number, track, incidence angle and heading angle of ascending and descending tracks are given in Table S1. We generate the ascending and descending interferograms using single look complex (SLC) products through the GMT5SAR code (Sandwell et al., 2011). We mostly follow the default procedure for processing and filtering. We use the amplitude image and a 1-arc-second SRTM digital elevation model (Farr & Kobrick, 2000) for co-registration and to produce the topographic phase correction. We generate the unwrapped interferometric phase for the ascending and descending acquisitions. We improve the coregistration through 1) geometric alignment on the basis of precise orbit (Sansosti et al., 2006), 2) de-ramping of SLC before interpolation of data (Miranda et al., 2015) and 3) mitigation of mis-registration on the basis of the spectral technique (Prats-Iraola et al., 2012). Azimuth coregistration is more difficult in TOPS mode acquisition than conventional strip map (De Zan & Monti Guarnieri, 2006). After high-quality co-registration, we remove the effect of topography from the SAR interferogram. In the first stage, we use a Gaussian filter with a wavelength of 200 m, and in the second a Goldstein filter is applied to the interferograms (Goldstein & Werner, 1998; Baran et al., 2003). We use SNAPHU (Chen & Zebker, 2002) to unwrap the interferogram with the threshold coherence of 0.15. Each subswath of interferometric SAR acquisition is processed individually and independently within its corresponding coor-

65 dinates. In the end, geocoding is applied to transfer the radar coordinate system to the
66 geographic coordinate system.

67 With the aim to define the data region to consider in the inversion, we have made
68 a simulation of the expected surface deformation signal for a single fault, assuming an
69 average dislocation. The deformation is shown in Figure S1. We take similar parame-
70 ters as those inverted from the geodetic Bayesian inversion approach (length:16 km, width:
71 2.8 km, dip: 40 degrees, strike: 194 degrees, depth: 6.6 km, average slip: 28 cm). We find
72 the signal to be confined to a region of $0.2^\circ \times 0.2^\circ$ and to decay quickly from its central
73 part, with a noise level of about 2 mm. This region limits the square of useful data since
74 at greater distances we cannot expect to have any signal in the data, but just add data
75 with noise or with a signal which has nothing to do with the earthquake. The topogra-
76 phy of the deformed region is flat, which is a favorable situation in relation to atmospheric
77 effects since they are correlated with topography. Nonetheless, we have tested also the
78 interferograms on a wider region ($0.6^\circ \times 0.6^\circ$) taking acquisitions before and after the
79 2015 Dajal earthquake, as shown in Figure S2. The figure clearly shows the presence of
80 atmospheric noise to the west of the Dajal earthquake deformation zone (Figure S2a-
81 e), although we had applied the Generic Atmospheric Correction Online Service (GACOS)for
82 InSAR (Yu et al., 2018) and removed atmospheric noise. We find significant noise on the
83 western side of the coseismic deformation zone in the interferograms of 10 November to
84 17 October 2015 (Figure S2a). This noise signal is absent in the descending interfero-
85 gram from 01 October to 18 November 2015 (Figure S2f), which demonstrates that the
86 western signal on LOS on the ascending interferogram is due to noise. Moreover, the noise
87 is lower in the ascending track for the 10 November to 04 December 2015 interferogram
88 (Figure S2e). This shows, the noise is probably due to local strong rains that affect the
89 area in fall and is stronger where the topography rises steeply, which is to the western
90 side of the deformation zone. We can clearly observe the region outside $0.2^\circ \times 0.2^\circ$ area
91 is noisy and can reach up to 15-20 cm. The inclusion of this wider area has a significant
92 impact on the results, as the noise level is higher than the signal at those distances. We
93 perform inversions for the wider region on the GACOS corrected data and find the pos-
94 terior probability density (PDF) is not well converging, as shown in (Figure S5). The
95 uncertainty on the inversion with the wider area is high. This confirms that our selected
96 smaller region (black rectangle in Figure S2) is quite feasible for inversions and avoids
97 any unnecessary contribution of noise in the results.

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1.2 Error Estimation

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Variance and covariance of the datasets are generally estimated experimentally to characterize the InSAR data errors that have occurred due to phase decorrelation. The InSAR errors arise mainly due to varying ionosphere and water vapor content in the upper atmosphere and lower atmosphere respectively (Hanssen et al., 1999). The error can also be encountered due to steep topographic variations at the site and anisotropic spatial variability (Knospe & Jonsson, 2010). We estimate the spatial variability of both ascending and descending interferograms using a semi-variogram (Wackernagel, 2003) by measuring the dissimilarity. We use the unbounded exponential function to compute dissimilarities with nugget, sill, and range variances for the ascending and descending interferograms separately (Table S2). Subsampled points used for Bayesian inversion for both interferograms are shown in Figure S3b, d respectively.

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1.3 InSAR Modeling

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We apply the geodetic Bayesian inversion (GBIS) approach (Bagnardi & Hooper, 2018) to the ascending and descending space-borne SAR interferograms covering the 2015 Dajal earthquake (Table S1). After estimating the experimental semivariogram by masking out the deformation zone, we use full-resolution InSAR data and then subsample both ascending (number of subsamples: 238) and descending (number of subsamples: 178) interferograms based on an adaptive quadtree gradient-based algorithm (Simons et al., 2002; Jonsson & Eklundh, 2002) using a threshold variance as given in Table S2. The large dataset is commonly subsampled in the Bayesian inversion to reduce the large computation time, and achieve enough information for a successful inversion (Bagnardi & Hooper, 2018). We prefer the gradient-based algorithm because the density of the samples is directly proportional to the displacement gradient and it recursively divides the LOS displacement into further four polygons each time, unless it achieves the selected threshold variance for ascending ($1.0 \times 10^{-2} mm^2$) and descending ($4.5 \times 10^{-3} mm^2$) interferograms (Figure S3a, c). We generate a kinematic synthetic model for a uniform rectangular dislocation source (Okada, 1992). The ascending and descending interferograms show the surface deformation of approximately 45 and 50 mm along their respective LOS direction (Figure S4a, b). Synthetic models (Figure S4c, d) with a single fault patch agree well with InSAR observed data, with average residuals of the order of average 2.3 mm for both interferograms (Figure S4e, f). We efficiently categorize the posterior probability density (PDF) of the ruptured fault geometry of the 2015 Dajal earthquake with automatic step size using the Metropolis-Hastings algorithm and the Monte Carlo Markov chain method (Tarantola, 2005; Fukuda & Johnson, 2008; Hastings, 1970;

133 Metropolis et al., 1953; Wang et al., 2017). We use 10^6 iterations to define posterior PDF,
 134 discarding the first 20,000 samples. We have used the epicentral location from seismic
 135 waveform modeling retrieved from the US Geological Survey (USGS, 2020), and ISC
 136 (International Seismological Centre) (Lentas et al., 2019) database as prior information.
 137 The inversion is done on both the wider area of $0.6^\circ \times 0.6^\circ$, which extends beyond the
 138 central earthquake deformation zone, and the smaller area of $0.2^\circ \times 0.2^\circ$ (Figure S2).
 139 As discussed in section 1.1 the wider area is significantly affected by the atmospheric noise
 140 accentuated westwards of the epicenter (Figure S2d). The posterior probability distri-
 141 bution for the wider area with 95% confidence intervals inversion results is given in Fig-
 142 ure S5, for the smaller area in Figure S6. The noise in the wider area propagates into
 143 a greater uncertainty level and misfit in the results, for which reason this solution is dis-
 144 carded. The final ruptured fault geometry of the 2015 Dajal earthquake with the 95%
 145 confidence interval is given by fault dip ($40^\circ \pm 12$), strike ($194^\circ \pm 6$), length (14.7 km
 146 ± 2.8), width ($2.9 \text{ km} \pm 1.2$), and depth ($6.5 \text{ km} \pm 1.2$) (Table S3). We use 10^6 iterations
 147 to define posterior PDF, discarding the first 20,000 samples (Figure S6). We have also
 148 used the epicentral location from seismic waveform modeling retrieved from the USGS
 149 NEIC database as prior information. The calculated source fault geometry is given in
 150 Table S3.

151 Text S2

152 2 Kinematic Inversions and Folding

153 2.1 Geometry of FBF and FPF models

154 In FBF and FPF models, the geometry of the active axial surfaces is obtained as-
 155 suming the conservation of layer thickness, length, and cross-sectional area of the incom-
 156 ing sediment in a balanced cross-section (Suppe, 1983; Suppe & Medwedeff, 1990). As-
 157 suming no cut-off angle between the incoming thrust sheet and the basal décollement
 158 for the FBF model (Figure 2a, see main text), the axial surface must bisect the décollement-
 159 ramp system, resulting in an angle of 70° clockwise for the deeper décollement, and 110°
 160 anti-clockwise for the shallower décollement (Figure 2b, see main text). For the FPF (Figure 2c,
 161 see main text), the axial surface 1 bisects the ramp-décollement with 70° , the axial sur-
 162 faces 3 and 4 bisect the wedge above the top of the fold at 55° and 70° , respectively, whereas
 163 axial surface 2 bisects the frontal fold with 55° (Figure 2d, see main text). We consider
 164 V_1 , V_3 , the long-term slip-rate above and below the fault (dark line), and V_2 is the long-
 165 term slip-rate along the active axial surface (Figure 2a,b, see main text). The angle be-
 166 tween V_1 and V_3 is 40° , V_1 and V_2 is 70° forms the closed hodograph shown as an inset
 167 in the (Figure 2b, see main text). The hodograph of the axial surface 2 is the same, but

168 with a reversed sense of slip rates (Sathiakumar et al., 2020). The motion along the two
 169 top-fold axial surfaces 3, and 4 is explained by the closed hodographs (Figure 2d, see main
 170 text). V_2 is the long-term motion of the rocks in wedge above the two top-fold axial sur-
 171 faces parallel to the passive axial surface – connects the dipping ramp and the two top-
 172 fold axial surfaces – at an angle of 82.5° from the horizontal. V_4 is -41.25° from the ax-
 173 ial surface 4, and $V_3 = 83.5^\circ$ from axial surface 3 forms the tow closed hodographs at
 174 axial surface 4, 3 respectively (Figure 2d, see main text). We also consider alternative
 175 models with the conjugate dip direction for the ramp with a similar structure but the
 176 opposite sense of motion.

177 **Text S3**

178 **3 Potency density of the Dajal earthquake**

179 The potency density — the average change of strain around the earthquake, i.e.,
 180 the stress drop divided by the rigidity of the country rocks — is 44 micro-strain, com-
 181 parable with that of the 2013 Mw 7.7 Balochistan, Pakistan earthquake, which had a
 182 similar centroid depth, and consistent with the general trend of potency density of thrust
 183 earthquakes worldwide (Nanjundiah et al., 2020)

Table S1: Features of Sentinel-1A interferometric pairs

Image pair (yy/mm/dd)	Track	Perpendicular baseline (m)	Temporal baseline (days)	Incidence angle ($^\circ$)	Heading platform ($^\circ$)
2015/10/17 – 2015/11/10	<i>ASC – T071</i>	174	24	36 – 39	-12.59
2015/10/01 – 2015/11/18	<i>DSC – T005</i>	40	48	36 – 39	-167.35

Table S2: Detail of interferogram errors calculated using semi-variogram

Track	Sill (mm) ²	Nugget (mm) ²	Range (km)	Threshold (mm) ²	Subsample (points)
<i>ASC – T071</i>	0.12	1.4×10^{-4}	9.75	1.0×10^{-2}	238
<i>DSC – T005</i>	0.038	2.3×10^{-3}	12.07	4.5×10^{-2}	178

Table S3: The estimated source fault parameters, inverting both ascending and descending interferograms along with their uncertainties by assuming single fault plane and ignoring fault-bends

Model	Lon (°)	Lat (°)	Str (°)	Dip (°)	Rake (°)	L (km)	W (km)	Dep (km)	Slip (m)	Mo ($10^{17} Nm$)	M_w
USGS	70.326	29.638	194	30	70	-	-	15.5	-	3.096	5.59
	-	-	182	47	68	-	-	11.0	-	1.792	5.44
GCMT	-	-	186	41	56	-	-	12.0	-	3.471	5.63
ISC	70.353	29.618	192	46	45	-	-	19.8	-	-	5.5
InSAR	$70.28 \pm$ 0.03	$29.66 \pm$ 0.02	$194 \pm$ 6	$40 \pm$ 12	$79 \pm$ 10	$14.7 \pm$ 2.8	$2.9 \pm$ 1.2	$6.5 \pm$ 1.2	$0.28 \pm$ 0.14	$3.94 \pm$ 4.0	$5.66 \pm$ 0.30

Table S4: The FPF and FBF models with varying dips at 7.5 km depth and their corresponding residuals. The 40° dip results consistently with the lowest RMS residuals.

Dip (°)	Forward-vergent FPF (mm)	Forward-vergent FBF (mm)	Backward-vergent FPF (mm)	Backward-vergent FBF (mm)
30	3.06	2.91	3.27	2.90
40	2.59	2.80	2.72	2.86
50	3.06	2.76	2.80	2.92

Table S5: AIC, RMS and Reduced-Chi-square analysis for forward- and backward-vergent models, depth = 7.5 km, dip = 40°. The reduced-chi-square statistic is a measure of the squared difference between the observed and modeled values, considering the degrees of freedom, and the sample size

Models	N	Np	RSS (mm^2)	RMS (mm)	AIC	Reduced-Chi Square
Forward- vergent FPF	17331	675	0.1246	2.59	2707.64	0.720
Forward- vergent FBF	17331	675	0.1401	2.80	2708.52	0.820
Backward- vergent FPF	17331	675	0.1285	2.72	2708.04	0.725
Backward- vergent FBF	17331	675	0.1429	2.86	2708.88	0.792

FBF: Fault Bend Fold

FPF: Fault Propagation Fold

N: InSAR data points

Np: Model parameters

RSS: Residual sum of square

RMS: Root mean square

AIC: Akaike Information Criterion

Table S6: AIC, RMS for forward- vergent FPF model with dip = 40°with, and varying depth.

Depth (km)	N	Np	RSS (mm^2)	RMS (mm)	AIC
6.0	17331	675	0.2009	3.40	2712.57
7.5	17331	675	0.1227	2.59	2707.64
9.0	17331	750	0.1095	2.51	3006.92
10.5	17331	900	0.1022	2.42	3606.59

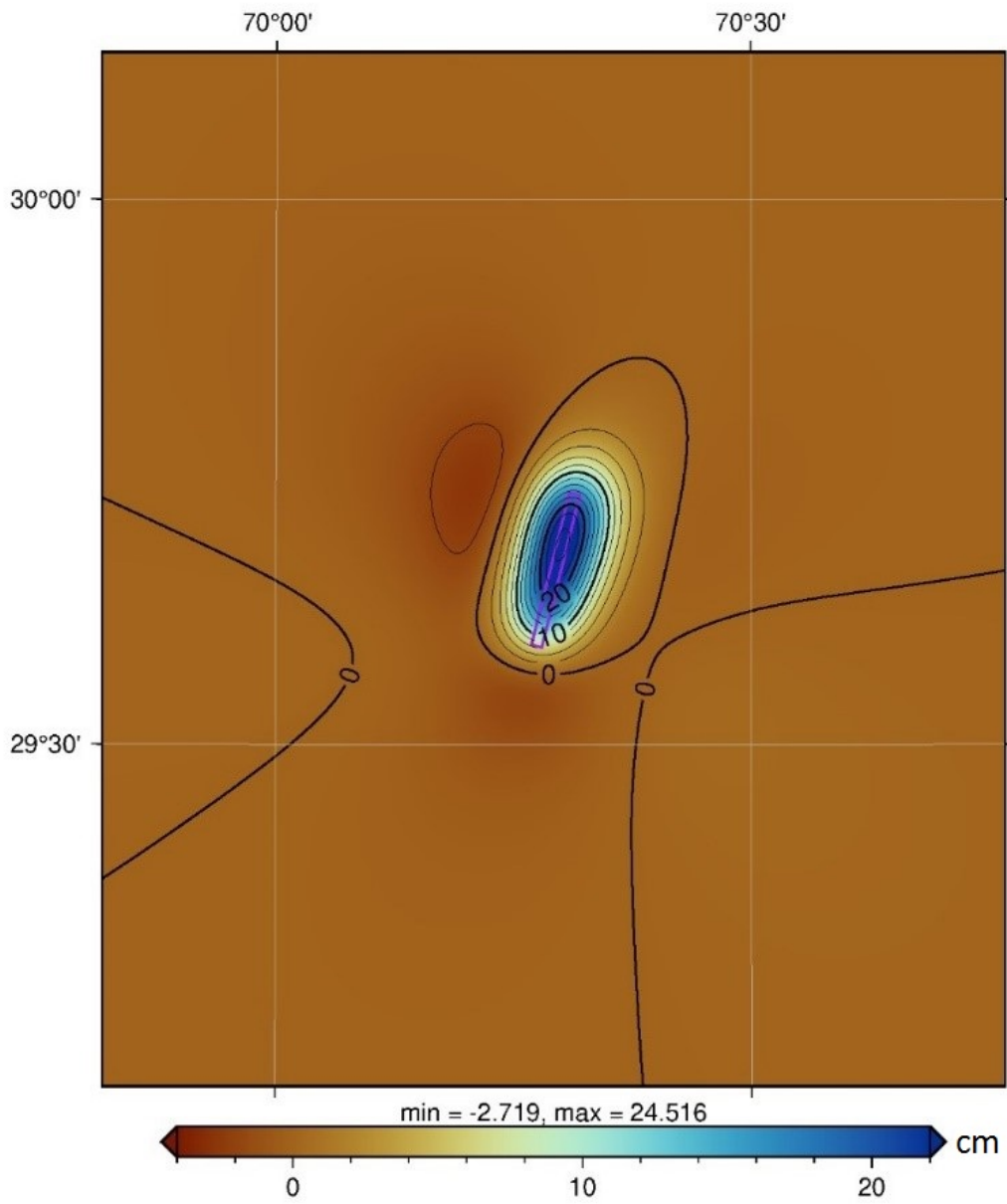


Figure S1: The simulation of the expected surface deformation signal for a single fault, assuming an average dislocation using the inverted fault parameters estimated from Bayesian inversion approach

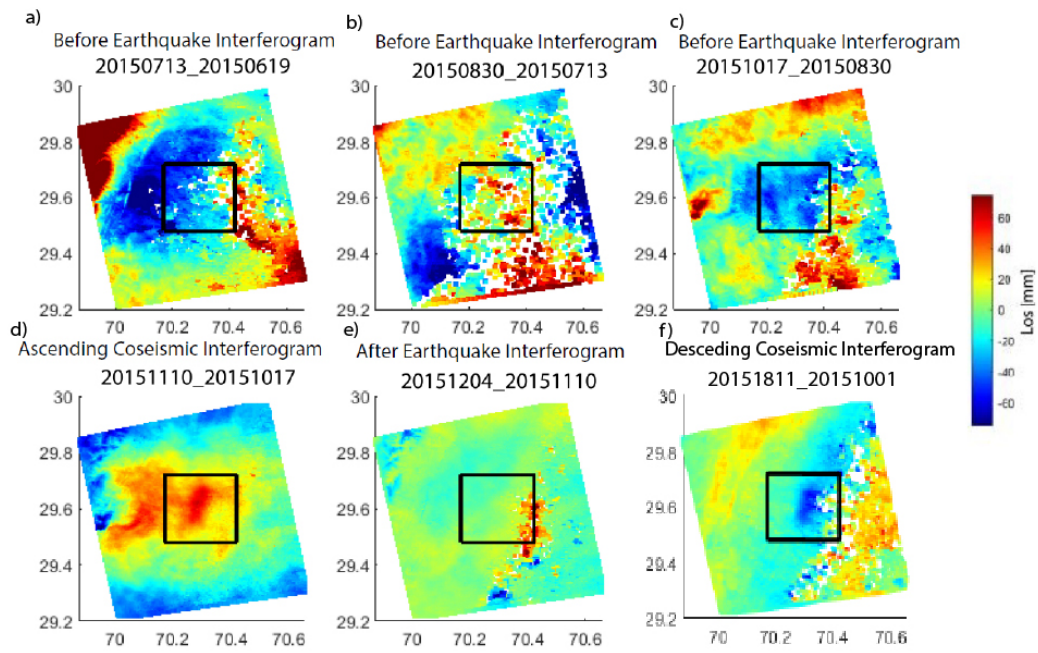


Figure S2: LOS displacement for ascending (a-e), and descending (f), calculated using Sentinel-1A interferograms for time couples before, after, and across the 2015 Dajal earthquake of 23 October 2015.

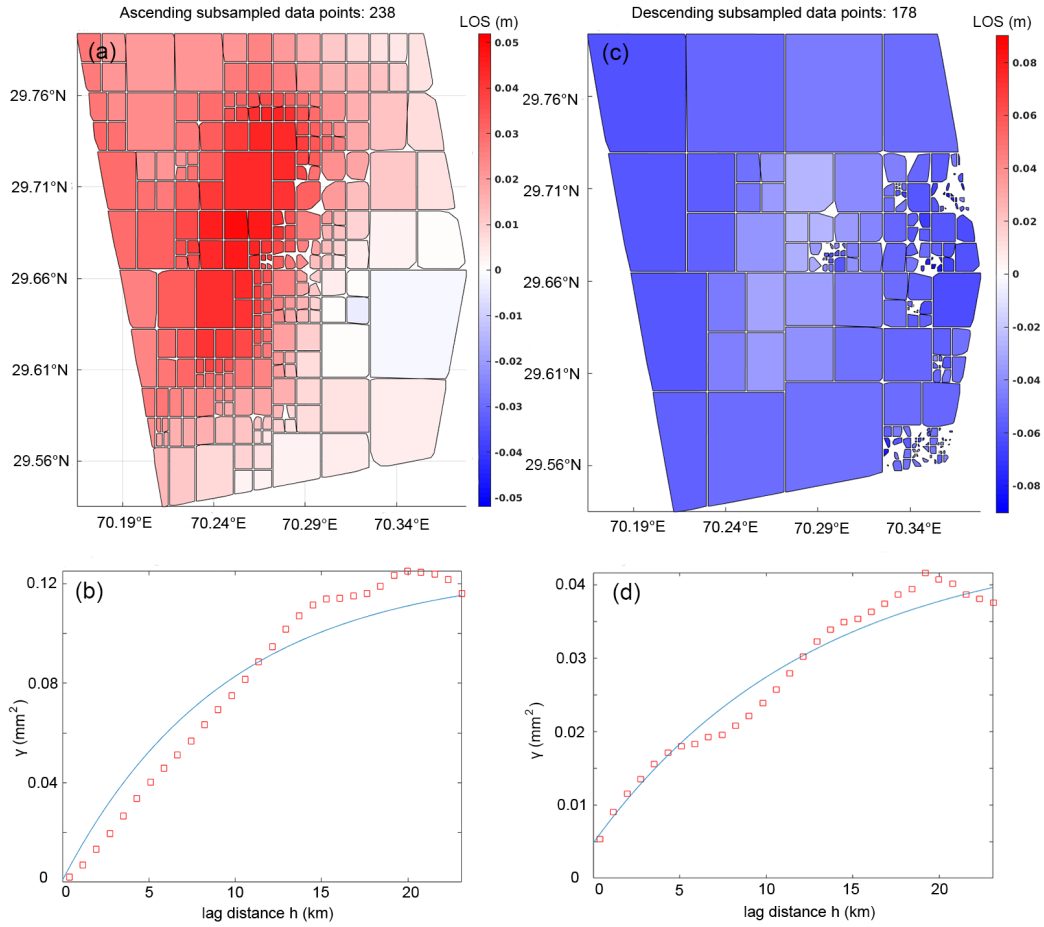


Figure S3: The adaptive quadtrees gradient-based subsampling and semi-variogram analysis. a, b) The subsampling and semi-variogram to estimate the covariance of the processed interferogram during ascending track respectively. The ascending interferogram is subdivided into 238 points with the threshold phase variance of $1.0 \times 10^{-2} \text{ mm}^2$. c, d) The subsampling and semi-variogram to estimate the covariance of the processed interferogram during descending track respectively. The descending interferogram is subdivided into 178 points with the threshold phase variance of $4.5 \times 10^{-2} \text{ mm}^2$. Blue (solid lines) is the exponential function of the semivariogram while Red (blocks) is the experimental semivariogram. The local origin for both interferograms is 70.289°N and 29.662°E .

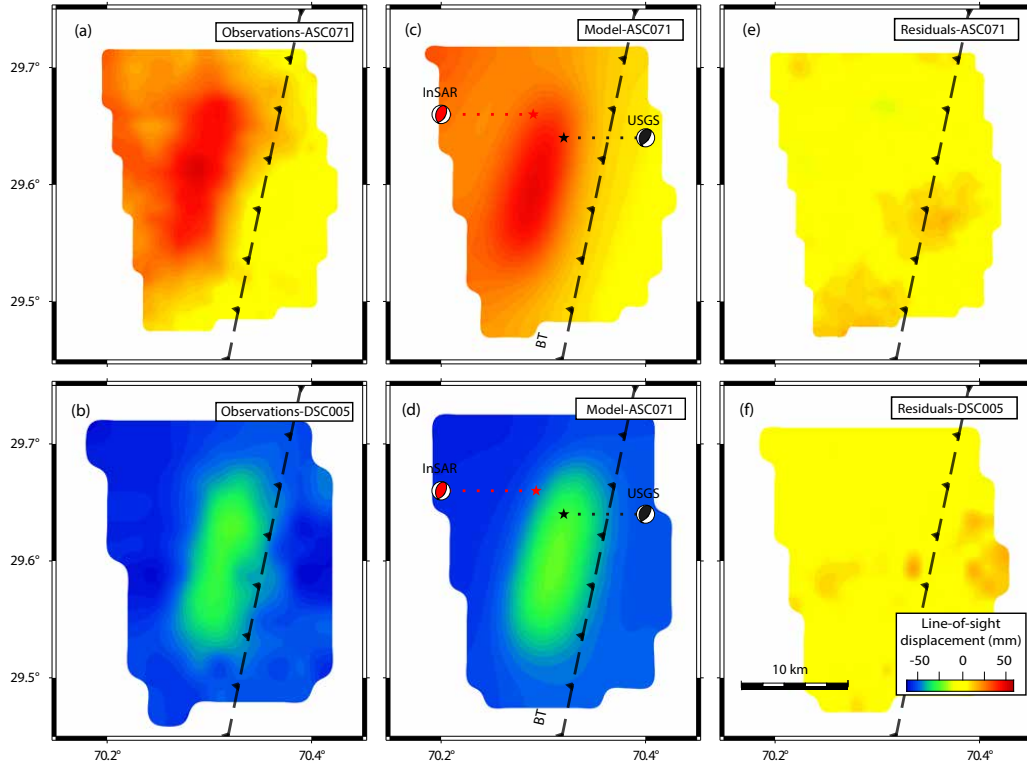


Figure S4: Single fault plane solution by using GBIS approach a, b) observation, c, d) synthetic interferogram and e, f) residual of the ascending track (T071) and descending track (T005) respectively. The deformation along the LOS displacement is approximately 50 mm and 45 mm along with less than 2.3 mm residual for descending and ascending track respectively. The focal mechanism solution (red colored) is produced by using inverted ruptured fault parameters and the focal mechanism solution (black colored) is taken from USGS.

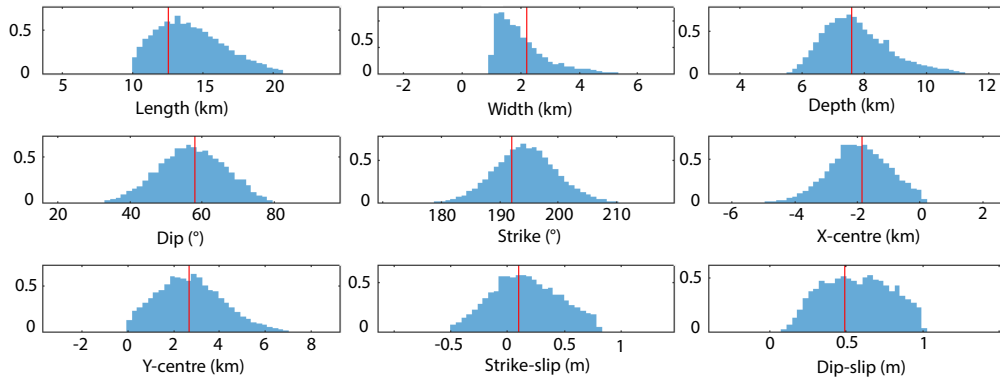


Figure S5: Histograms of source fault model parameters through Bayesian inversion approach with 10^6 samples, on wider area of $0.6^\circ \times 0.6^\circ$. The y-axis represents the probability density, and the red line indicates the optimal model values with 95% confidence interval for GACOS corrected data. (Rejected solution).

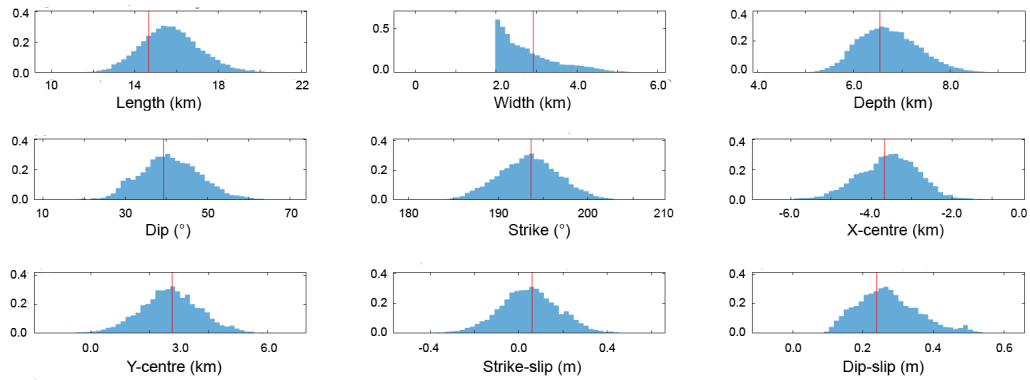


Figure S6: Histograms of source fault model parameters through Bayesian inversion approach with 10^6 samples, on area of $0.2^\circ \times 0.2^\circ$. The y-axis represents the probability density, and the red line indicates the optimal model values with 95% confidence interval. The optimal model has fault length around 15 km, width 2.9 km, depth 6.5 km, dip 40° , strike 194° , and average slip 0.28 m. (Best solution).

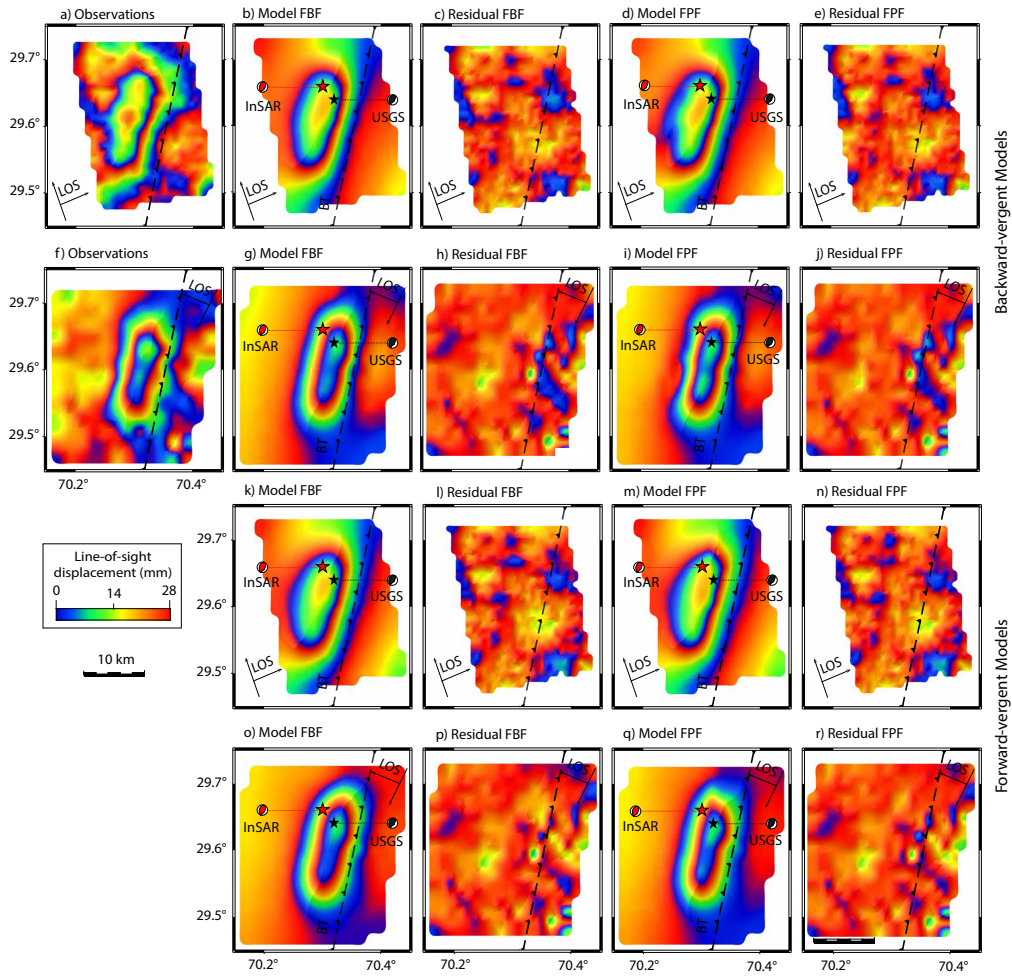


Figure S7: The FBF, and FPF model solutions a, f) ascending and descending InSAR observations, b - i) Ascending and descending backward-vergent fault-bend fold and fault-propagation fold models and residuals, k - r) Ascending and descending forward-vergent fault-bend fold and fault-propagation fold models and residuals. The maximum co-seismic slip along the LOS displacement is approximately 50 mm along with less than 2.8 mm residual for descending and ascending track respectively. Focal mechanism solution (red colored) is produced by using inverted ruptured fault parameters and focal mechanism solution (black colored) is taken from USGS.

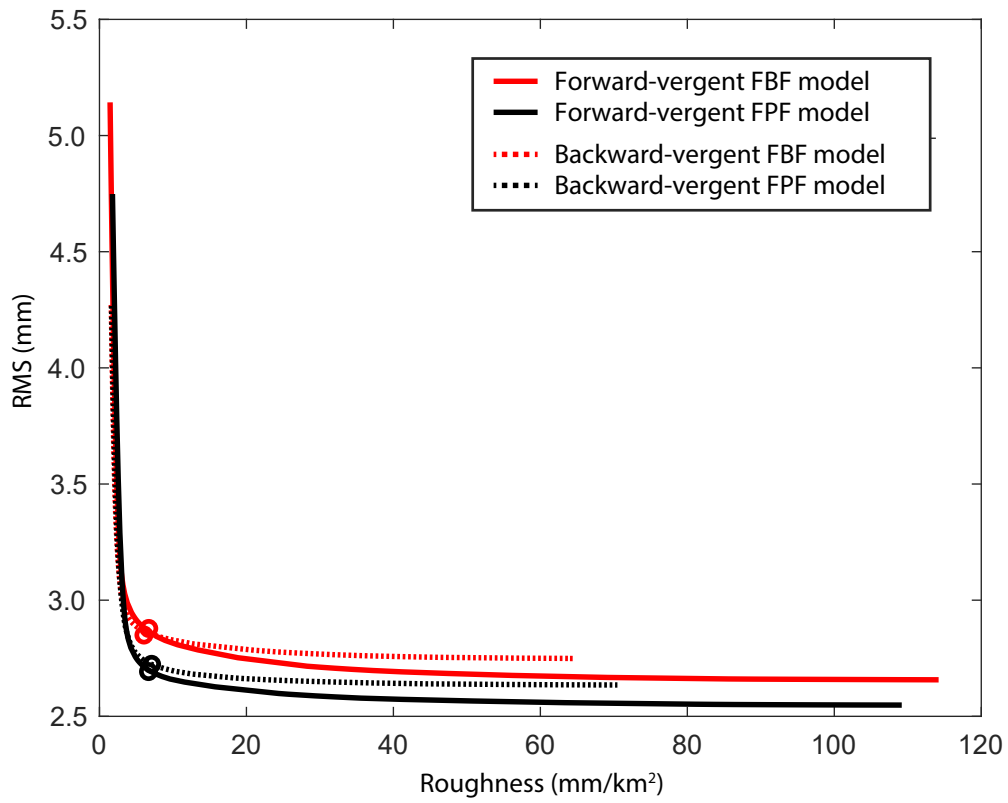


Figure S8: L-curve of RMS for forward- and backward-vergent FPF and FBF models. FPF: Fault-propagation fold, FBF: fault-bend fold

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